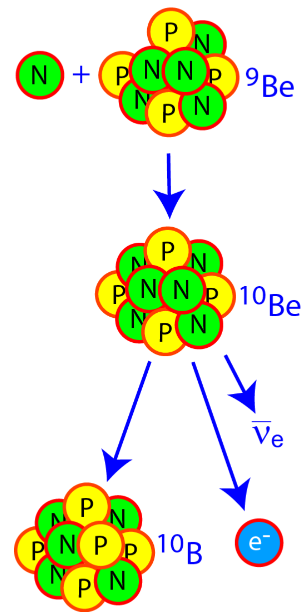
Neutron Capture Processes

Many elements are created by fusion in stars, but heavy elements can't be! Those elements may result from neutron capture processes:

1. Free neutrons are created by nuclear reactions in a red giant star.
2. A stable nucleus in the star (Be-9 in the example at right) absorbs a neutron, making a neutron-rich and unstable Be-10.
3. The Be-10 nucleus releases energy/becomes stable by beta decay, turning a neutron into a proton and forming B-10.
4. Thus, a Beryllium nucleus has been turned into a heavier element, Boron!
5. This new stable Boron nucleus might absorb a neutron, and the whole process continues.

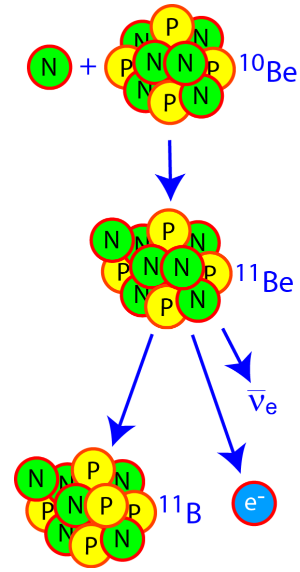


Capture or decay first?

In a star with loose neutrons around, stable nuclei might absorb them every so often. Unstable nuclei might too, as long as they didn't decay first. Consider the Be-10 above... its half-life is over 1 million years. *If neutrons were abundant enough that a nucleus would normally capture one every ten years, would Be-10 be more likely to capture or decay first?*

A Be-10 nucleus capturing a neutron would become Be-11, which has a half-life of 13.8 seconds. *With the same assumptions above, would it be more likely to capture or decay first?*

The number of neutrons available and the half-life of each isotope determines whether it is more likely to capture a neutron or decay!



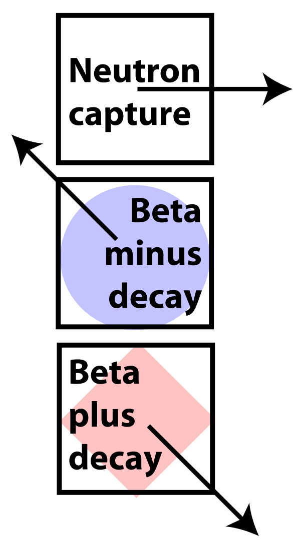
Creating a model

Note the "legend" at right: on a chart of the nuclides, neutron capture moves a nucleus to the right, while beta decays go up & left or down & right. **Remember this for the next part!**

Let's construct a simple model of how neutron capture occurs in a red giant star. In our model:

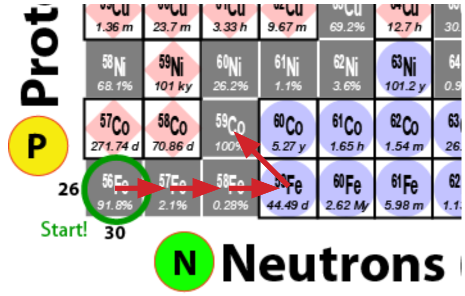
1. Neutrons capture every 10 yr
2. Isotopes with half-lives longer than 10 yr will capture, while isotopes with half-lives shorter than 10 yr will decay.

Using that model, you can make a prediction of the "s-process", a set of slow neutron capture reactions that may occur in a red giant star!



S-process simulation

Use your own “Neutron Capture Processes” Chart, but work with a partner to share ideas:



1. *Assume a starting nucleus of Fe-56, and that neutron capture occurs **every 10 years**.*
2. *Decide whether Fe-56 will capture or decay first (hint: it’s stable, so the half-life is forever) & draw an arrow to the resulting nucleus*
3. *Decide whether the resulting nucleus will decay or capture first & draw an arrow; repeat!*
4. *Continue until you make the heaviest Sr (strontium) isotope you can.*

Creating new isotopes

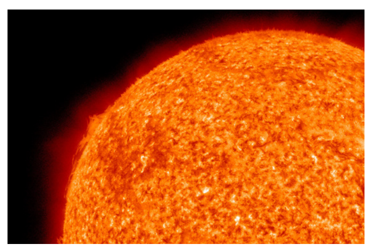
Did your s-process calculation create nuclei that were close to the line of stable (grey) isotopes, or far from it?

When the s-process ends, all unstable nuclei will decay back to stability (generally beta-minus, up and to the left). *If you have time, draw arrows (in a different color) from the unstable isotopes you created to the stable isotopes they will become.*

Remember, we’re trying to understand if the s-process can create the heavy elements we see in nature. *Were all stable isotopes in this part of the chart created during the s-process (or after its unstable isotopes decay)?*

R-process simulation

According to your s-process simulation, red giants may be able to make many heavy elements. However, it didn’t produce all the heavy stable isotopes, so it can’t explain why they are found in nature. There may be other nuclear processes responsible!



In a supernova or neutron star merger, the density of free neutrons is likely much higher than in a red giant star. In those environments, neutron capture would proceed much faster! This is called the rapid-neutron-capture or “r-process”.

Use your own Chart, but work with a partner to share ideas:

1. *Assume a starting nucleus of Fe-56, and neutron capture **every 100 ms**.*
2. *Again, for your starting nucleus and each one you make, decide whether it will capture or decay first & draw an arrow to the resulting nucleus*
3. *Continue until you make the heaviest Sr (strontium) isotope you can.*

When the r-process ends, all unstable nuclei will decay back to stability (generally beta-minus, up and to the left). *If you have time, draw arrows (in a different color) from the unstable isotopes you created to the stable isotopes they will become.*

Maybe the r-process can also create the heavy elements we see in nature. *Which stable isotopes could be made by the r-process or s-process? Which could only come from the r-process? Which were only made by the s-process?*